

Deep learning denoising by dimension reduction Application to the ORION-B line cubes

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1/ Hyperspectral line cubes

Large bandwidth receivers for millimeter radio-telescopes allows the acquisition of position-position-velocity (x, y, v) data cubes of the interstellar medium over a wide field of view and a broad frequency coverage. [1]
Due to the Doppler effect of the moving gas, these lines do not have a dirac-shaped emission spectrum, but rather an envelope with a Gaussian appearance.
Knowing the emission frequency of a molecule, we can interpret the spectral bands as an image of the emission a given molecule moving at a certain radial velocity.

3/ Informed local autoencoder

- \blacktriangleright In order to encompass those challenges, we propose significant improvements of typical autoencoder neural network.
- 1. As distant channels are independent, they should not be encoded together. We



• These cubes contain much information on the physical and chemical properties of the emitting gas but their large size coupled with inhomogenous signal-to-noise ratio are major challenges for consistent analysis and interpretation. propose an alternative autoencoder architecture.



Dense and locally dense layers

Example of locally dense autoencoder

The maximum distance between two channels encoded together is defined from the data analysis as the characteristic distance between two channels at which the remaining mutual information is of 5%.

2. Moreover, as the major part of the voxels (samples (x, y, v)) are signal free, we want the autoencoder to encourage the sparcity of the denoised cube.





CO isotopologues emission (9 km/s) Ex. of noise map

Goals

- \star Increase the signal-to-noise ratio, in particular in the region where it is low.
- \star Reduce the dimension of the cube over the spectral axis.
- \rightarrow An autoencoder neural network could do both. [2]

2/ Challenges

• As distant channels are statistically independant, the redundancy that can be exploited is only local.





where σ_k is the noise RMS for the k-th pixel and w_{jk} and priori from another segmentation method indicating whether a pixel is likely to contain signal or not.

4/ Results

• We apply this method on some cubes. For each one, a new network has to be trained.



Example of noisy and denoised channel and the residues for ^{13}CO (1-0).

- We compare the denoising with that of the Gaussian fit method ROHSA. [3]
- We observe different behaviors, in particular the Gaussian fit denoise whatever the signal level and tends to distort the signal of high intensity. The autoencoder tends to preserve these regimes.
- The distortions of the signal (regions where



 13 CO (1-0) emission, 7 and 10 km/s channels

→ The information shared with a given channel decreases with the distance. As a consequence, the amount of exploitable mutual information (MI) is very limited.



Band pairs MI for ¹³CO (1-0) and the Indian Pines dataset

the residuals are correlated) are quite low for the autoencoder.

5/ Conclusion

Statistical denoising with low redundancy
 Data analysis based architecture and loss function
 Denoising of the low level signal
 No significant dimension reduction



6/ **References**

Pety et al., The anatomy of the Orion B giant molecular cloud [...], 2017
 Licciardi et al., Nonlinear PCA for visible and thermal hyperspectral images [...], 2015
 Marchal et al., ROHSA: Regularized Optimization for Hyper-Spectral Analysis-Application [...], 2019